

Why go to space?

Why go to space?

To explore.

- Colonization, tourism, and the expansion of the limits of humanity
- Science (astrobiology, geology, physics, etc.)

For resources.

- Rare elements (gold, platinum, etc.)
- In-situ resource utilization
- Water
- Data

To look out.

- Astrophysics
- Atmospheric sounding

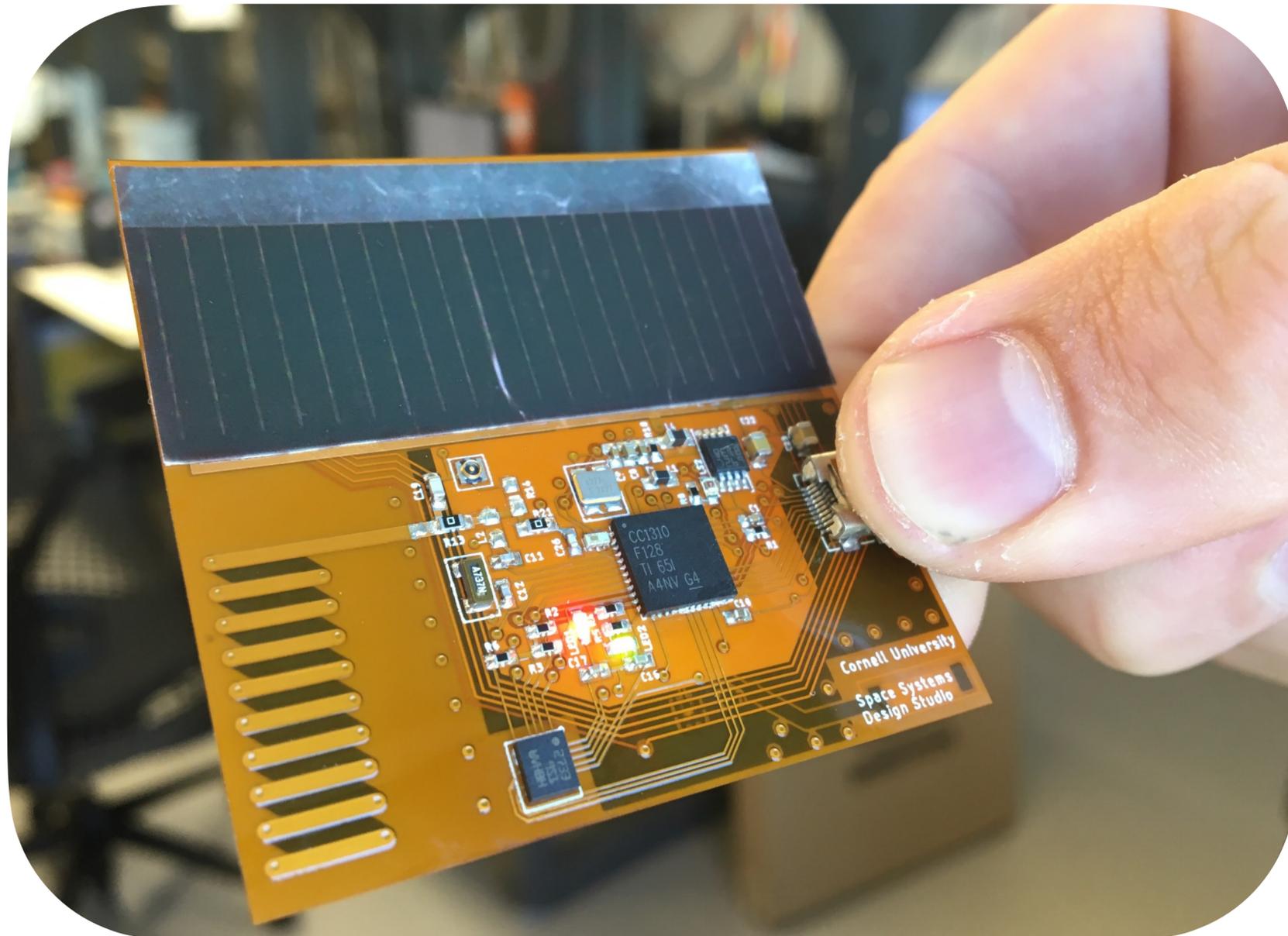
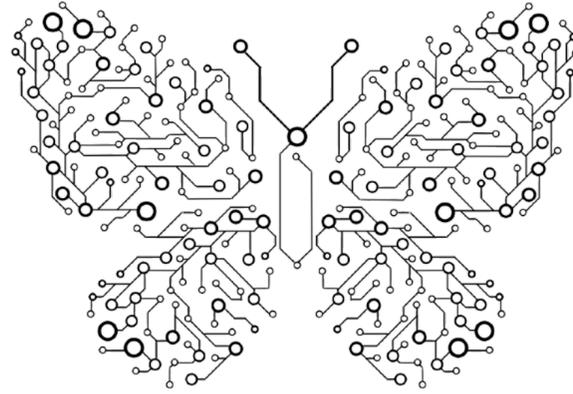
To look down.

- Surveillance (military and commercial)
- Weather forecasting/science

For microgravity.

- Science and materials processing

Monarch



I'm Hunter Adams

- Received a PhD in aerospace engineering (all space, no air).
- My research surrounds (very) small spacecraft called “chipsats.”

Going to space is hard and expensive.

- Getting permission to go to space is hard . . .
- Traveling to space is hard . . .
- Being in space is hard . . .

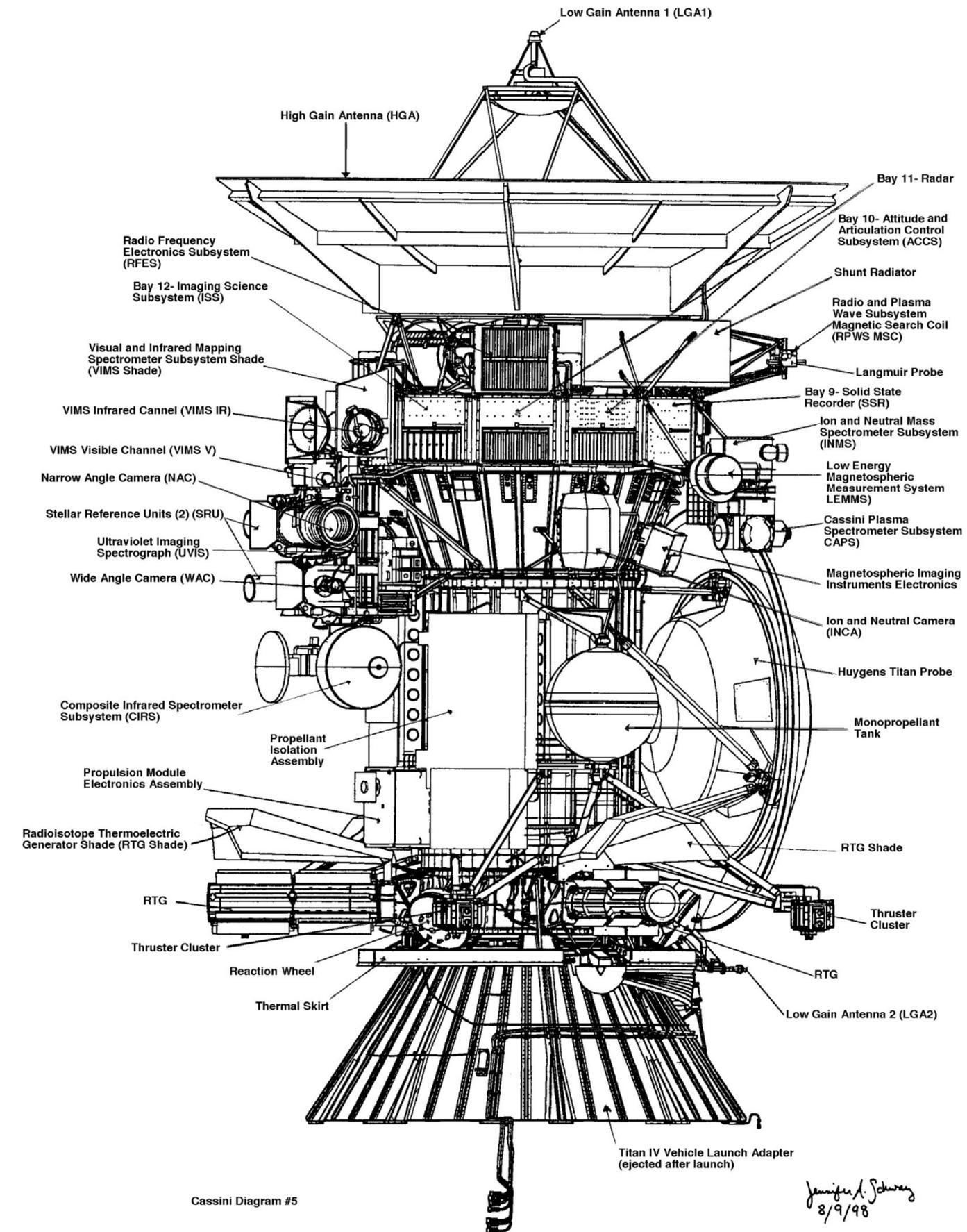
- Getting permission to go to space is hard . . .
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. . . in ways that are unique enough to require separate consideration.

We think about complex systems
in terms of subsystems.

Spacecraft subsystems

- Attitude determination and control
- Propulsion
- Telemetry and command
- Avionics
- Flight software
- Power
- Thermal
- Structures
- Entry, descent, and landing
- Environmental control and life support systems



This course exists to understand spacecraft at a systems-level, and to understand the environment in which spacecraft exist.

Course goals

- To understand space missions and systems, and how the space environment and mission requirements drive spacecraft design.
- To understand the fundamentals of spacecraft subsystems, including propulsion, attitude determination and control, power, structures, thermal, communications, and command and data handling.
- To understand typical practices for designing space systems in a contemporary context of US commercial space and government agencies.
- To simulate a spacecraft in operation at the level of a Preliminary Design Review (PDR) using state of the art tools, and identify and characterize subsystems for a preliminary spacecraft design.

Today's goal

To introduce the topics of this course via a case study.

Cassini

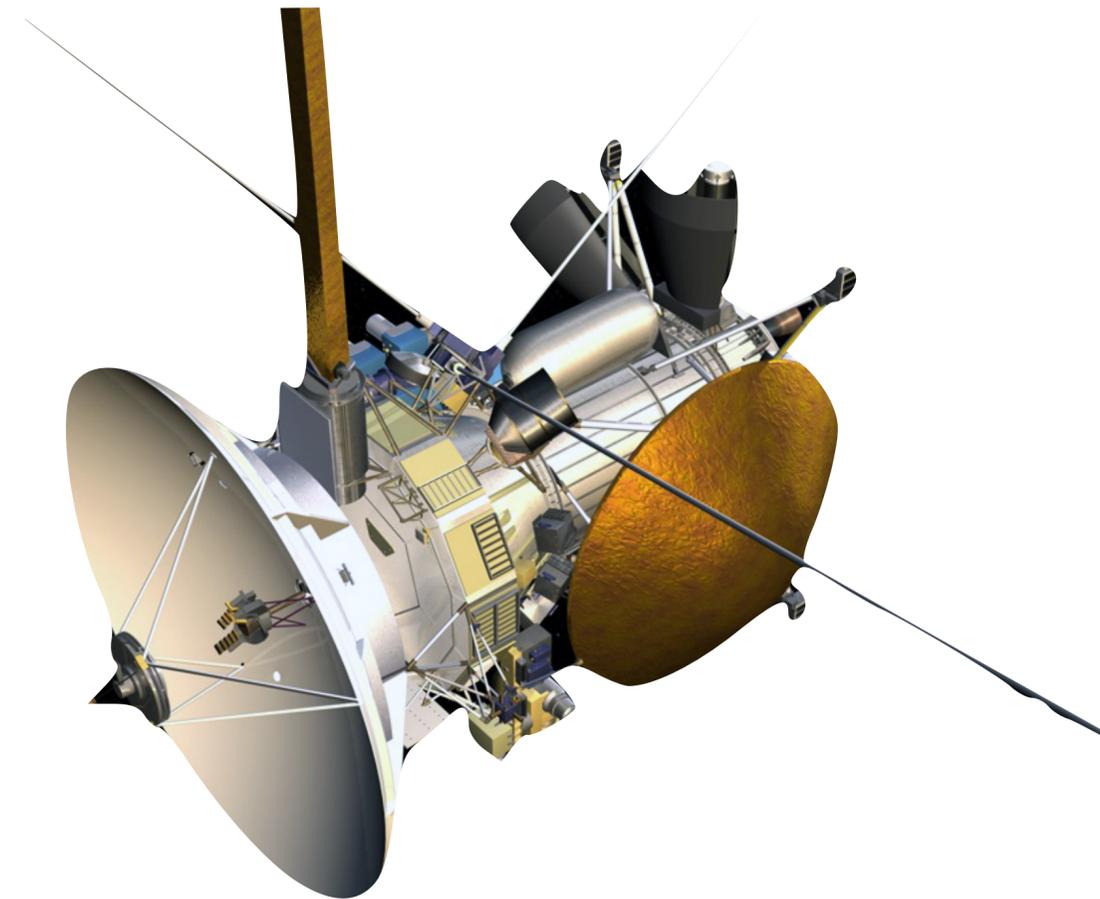


Why start with a case study?

To avoid the "wait, what am I even doing?" moment.

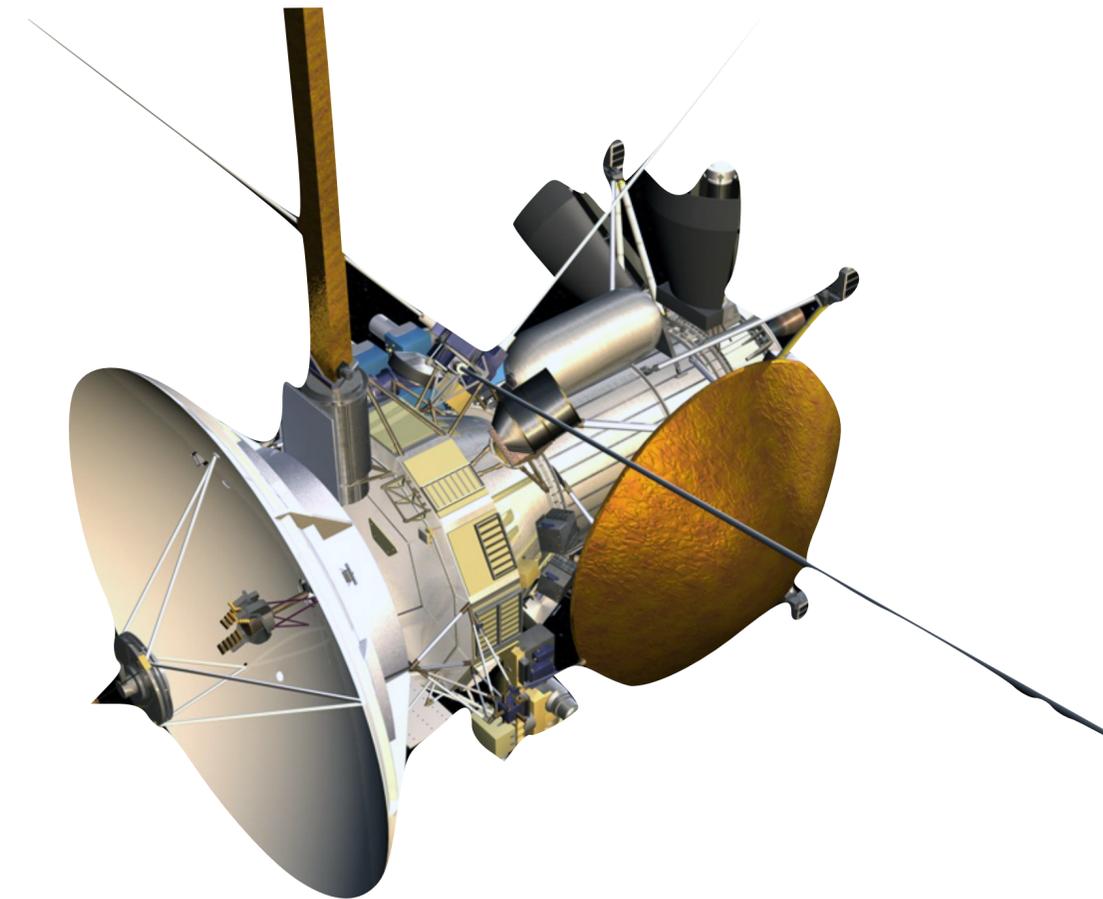
Why feature Cassini?

- It is familiar.
- It was radically successful.
 - Oct. 15, 1997 - Sept. 15, 2017
 - Gathered 635 GB of data
 - Discovered 6 moons
 - Took over 450,000 photographs
 - Deployed Huygens Probe to Titan
 - Led to the publication of >4000 papers



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*Cassini was a success because of its **design**, and because of the **process by which it was designed**.*

That process begins with *objectives*:

- Defines the *ends* to which the spacecraft is the *means*.
- May be fuzzy and imprecise.

SATURN

- Determine temperature field, cloud properties and composition of the atmosphere of Saturn.
- Measure the global wind field, including wave and Eddy components; Observe synoptic cloud features and processes.
- Infer the internal structure and rotation of the deep atmosphere.
- Study the diurnal variations and magnetic control of the ionosphere of Saturn.
- Provide observational constraints (gas composition, isotope ratios, heat flux,...) on scenarios for the formation and the evolution of Saturn.
- Investigate the sources and the morphology of Saturn lightning (Saturn Electrostatic Discharges (SED), lightning whistlers).

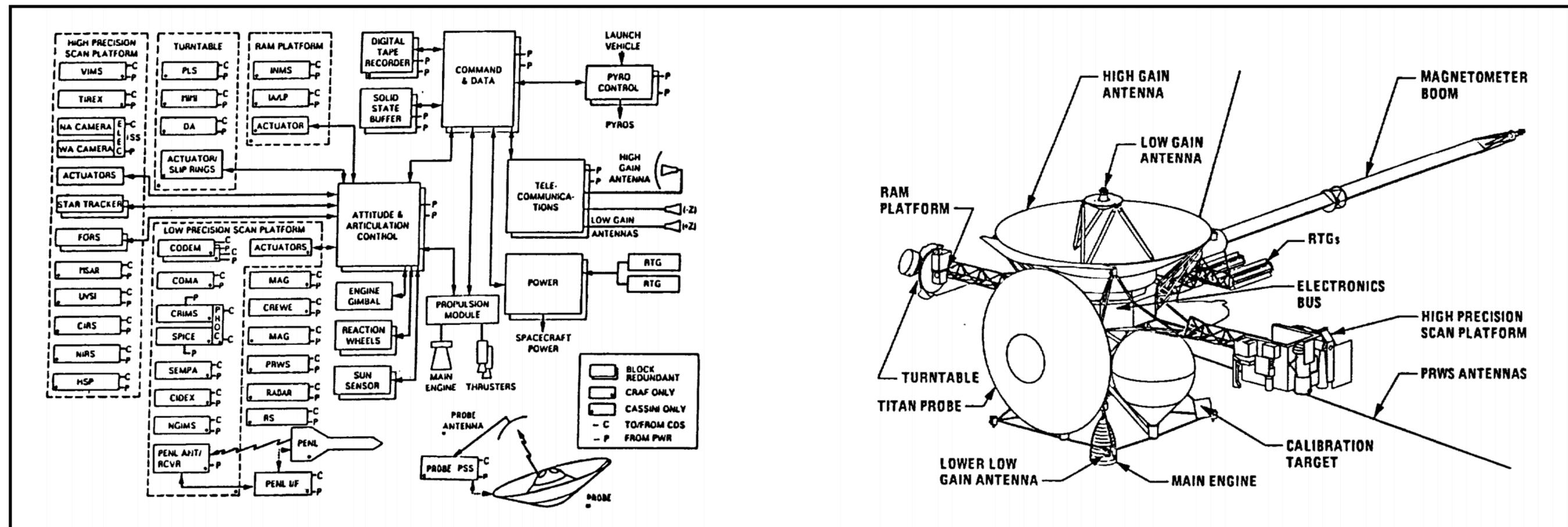
Objectives become requirements:

- Unambiguous, isolated, concise, measurable, unique, and consistent.
- "What must the spacecraft do?"
- Analogous to test-driven software development

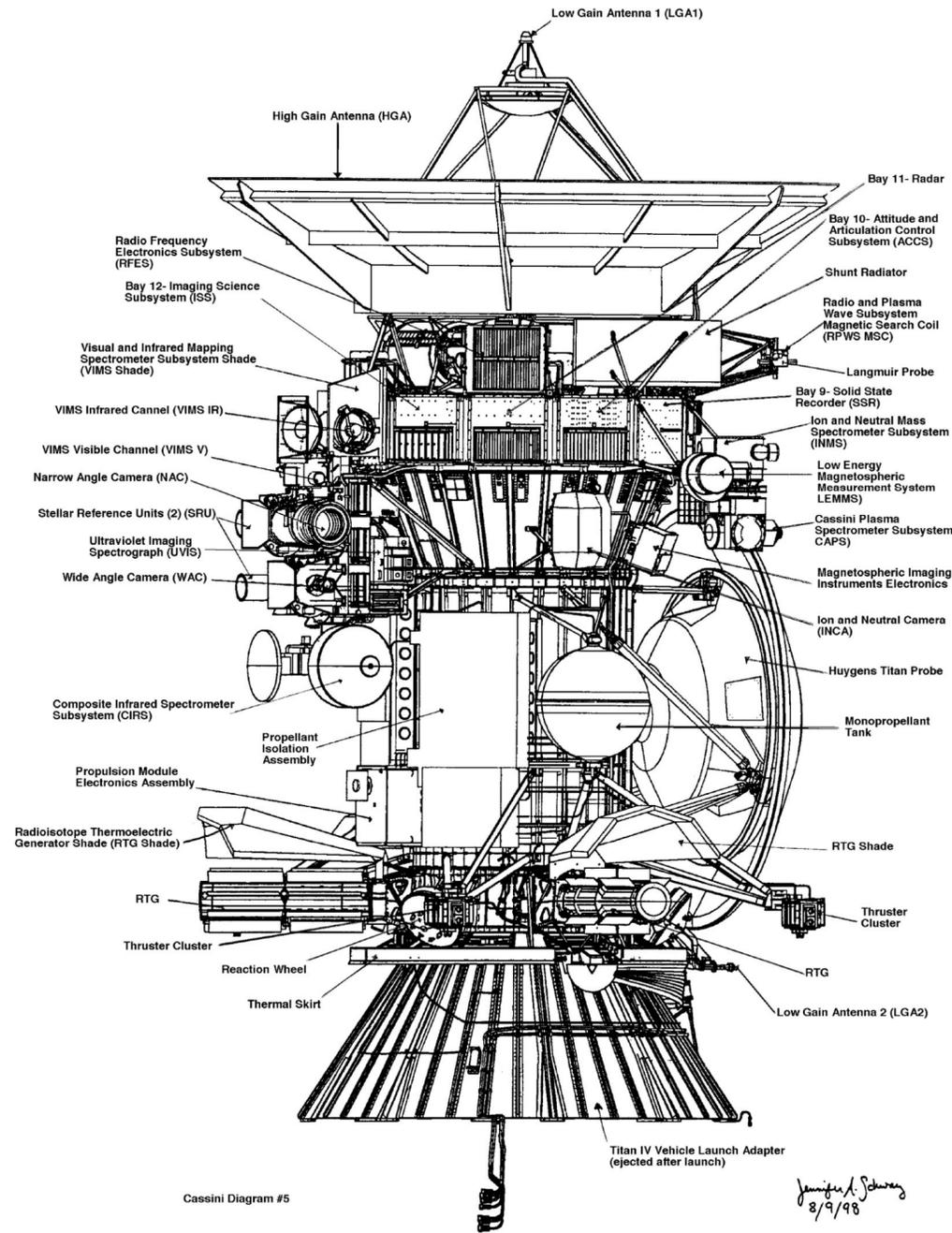
- The Orbiter shall provide telecommunications to and from 34 m and 70 m DSN stations at multiple data rates up to 115.2 kbps at the asteroid and 67.9 kbps at Saturn, and shall provide on-board storage for at least 3.6×10^9 bits of data for use during cruise, Probe mission and the tour.
- The Orbiter shall provide adequate pointing accuracy at Probe-Orbiter separation to establish the proper link geometry and the capability to receive Probe data for the full range of relay link geometries.
- The Orbiter shall be designed to withstand 52 R_J Jupiter flyby, Saturnian and ring plane crossing environments in the clear zones, sparsely populated regions and upper fringes of Titan's atmosphere.

Requirements become design:

- System requirements flowdown to subsystem requirements
- Subsystems are designed to satisfy requirements.
- Design is complete when all requirements are met.



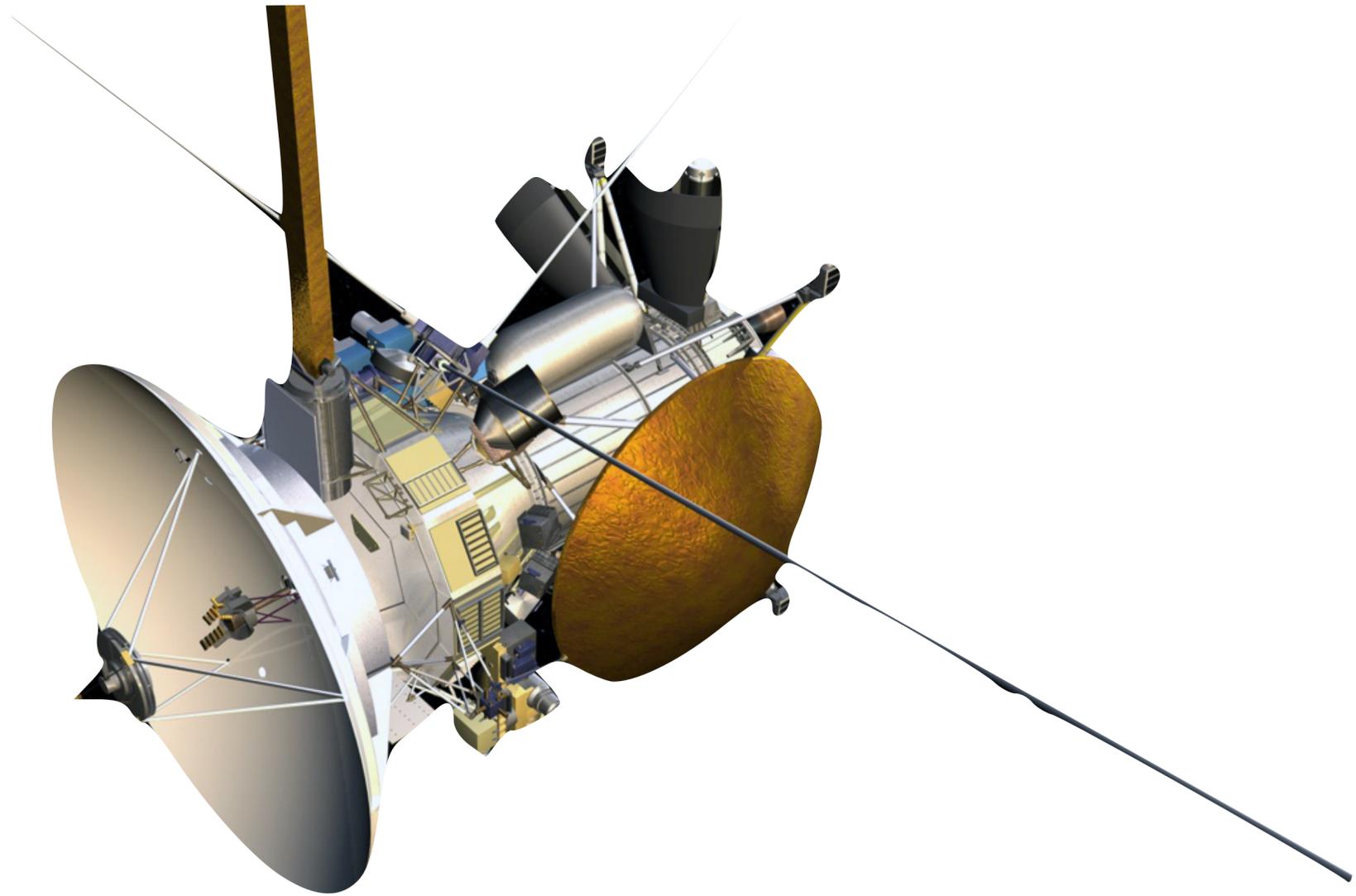
Design is finalized, and the spacecraft is *fabricated*



What did they design?

Let's consider Cassini one subsystem at a time.

Power

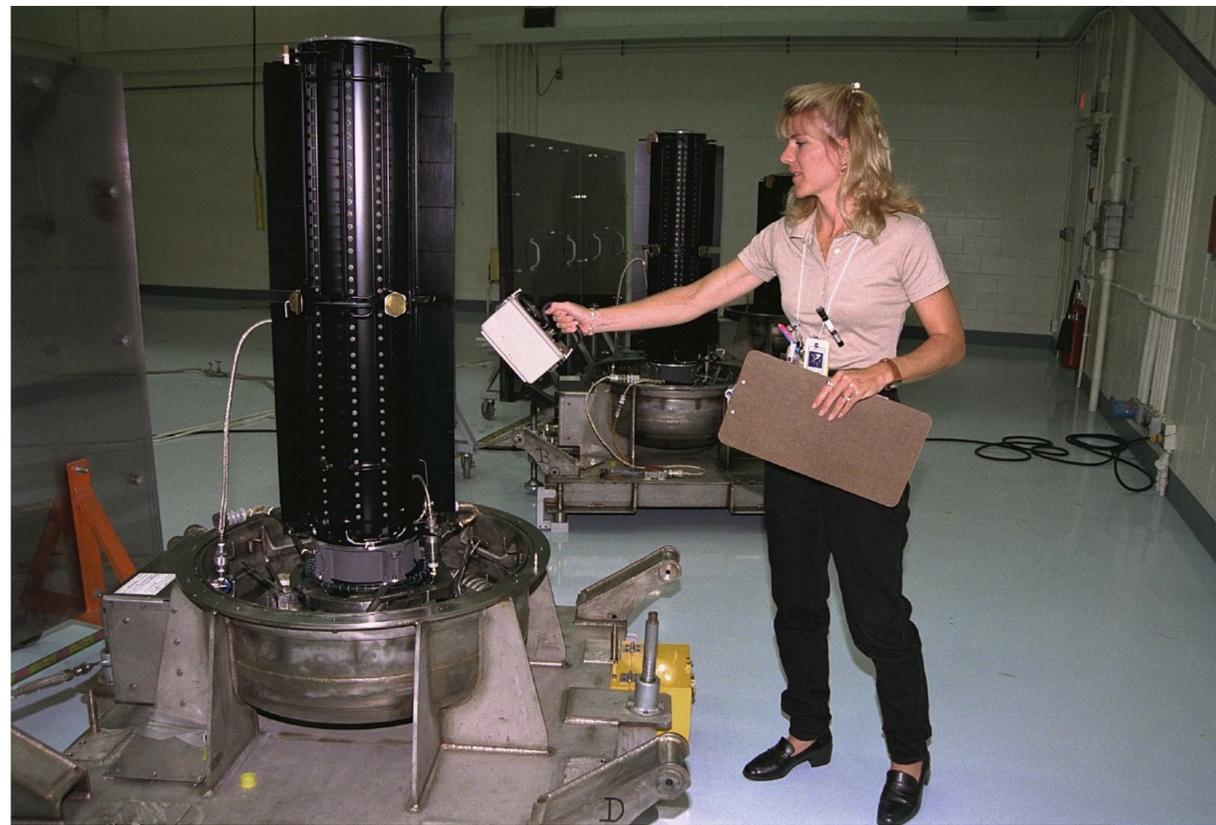


What's missing?

Power

Radioisotope Thermoelectric Generator (RTG)

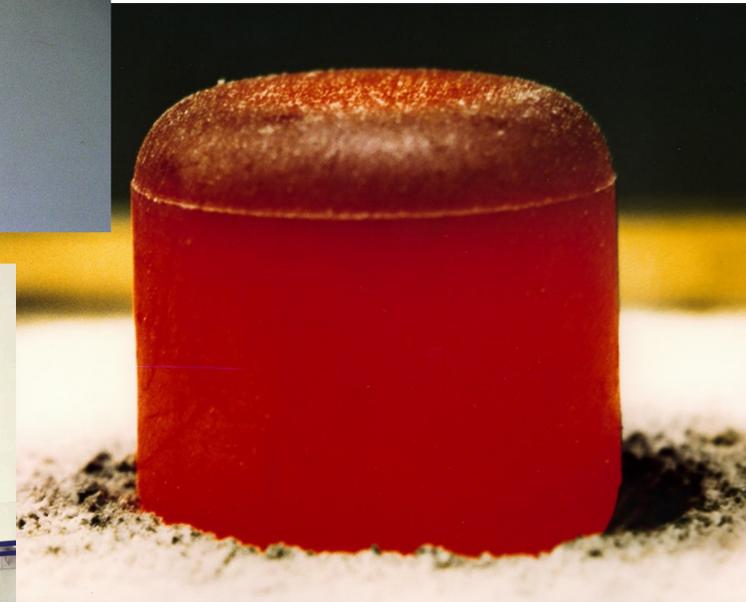
- Cassini carried a 33 kg slug of Plutonium-238
- Converted heat generated by radioactive decay into power
- A lot of power. At the end of Cassini's life, the RTG's were still continuously generating 600-700W
- Also carried NiCd batteries for storage and power peaks



Cassini's RTG



RTG on *New Horizons*



Cassini's Plutonium

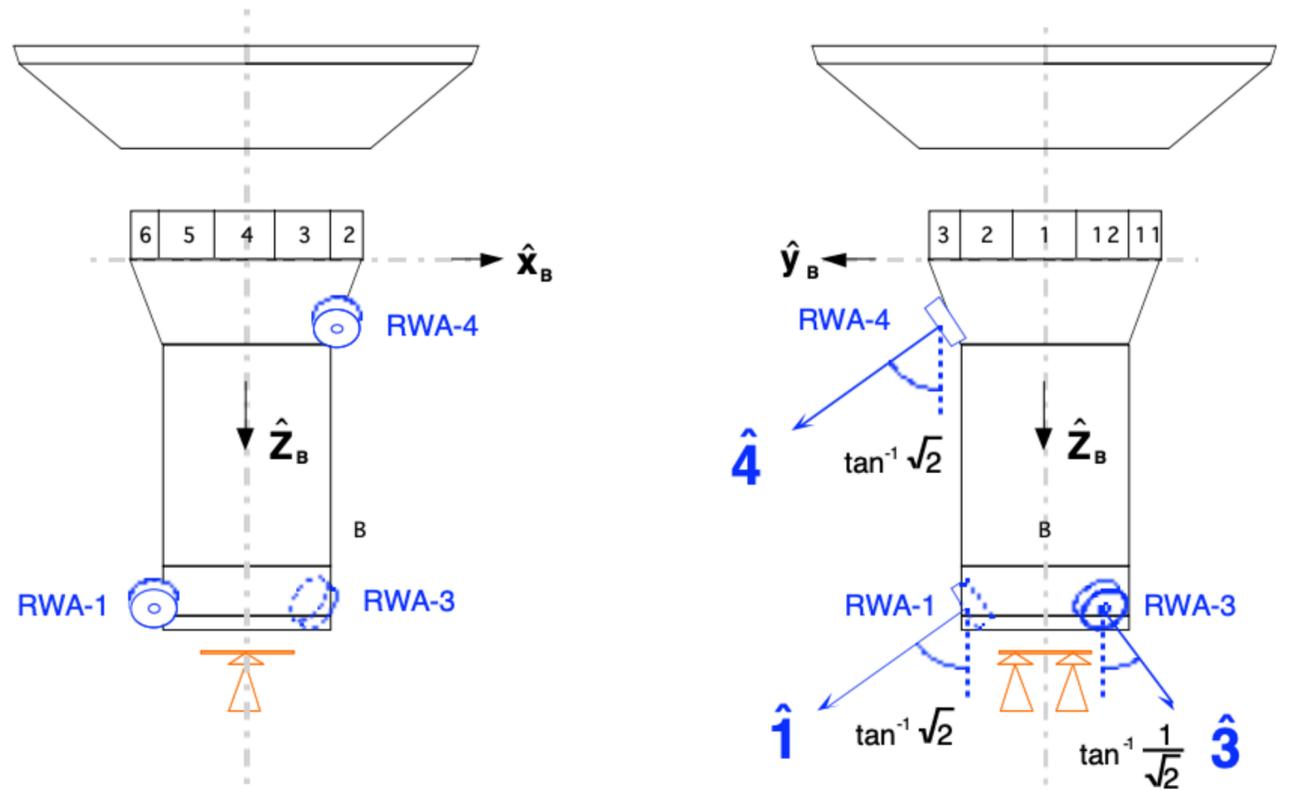
Attitude Control

Accuracy Requirements	Requirements
HGA pointing control requirement (radial 99%)	
X-band (Telecommunications)	3.2 mrad
Ka-band (Radio Science)	2.0 mrad
Ku-band (Radar Mapping of Titan)	4.6 mrad
S-band (Huygens Probe Relay)	6.0 mrad
LGA X-band pointing control requirement (radial 99%)	4.0 degrees
Science inertial pointing requirements (radial 99%)	
Control	2.0 mrad
Knowledge	1.0 mrad
Science pointing stability requirements (2σ per axis) for time windows of: ⁹	
0.5 s	4 μ rad
1 s	8 μ rad
5 s	36 μ rad
22 s	100 μ rad
100 s	160 μ rad
900 s	200 μ rad
1200 s	220 μ rad
1 hour	280 μ rad
Main engine ΔV burns (1σ):	
Fixed Magnitude	10 mm/s
Fixed Pointing	17.5 mm/s
Proportional Magnitude	0.2 %
Proportional Pointing	3.5 mrad
Thruster ΔV burns (1σ):	
Fixed Magnitude	3.5 mm/s
Fixed Pointing	3.5 mm/s
Proportional Magnitude	2 %
Proportional Pointing	12 mrad

Do you think Cassini was
3-axis stabilized?

What sorts of actuators do
you think Cassini used to
meet these requirements?

Attitude Control



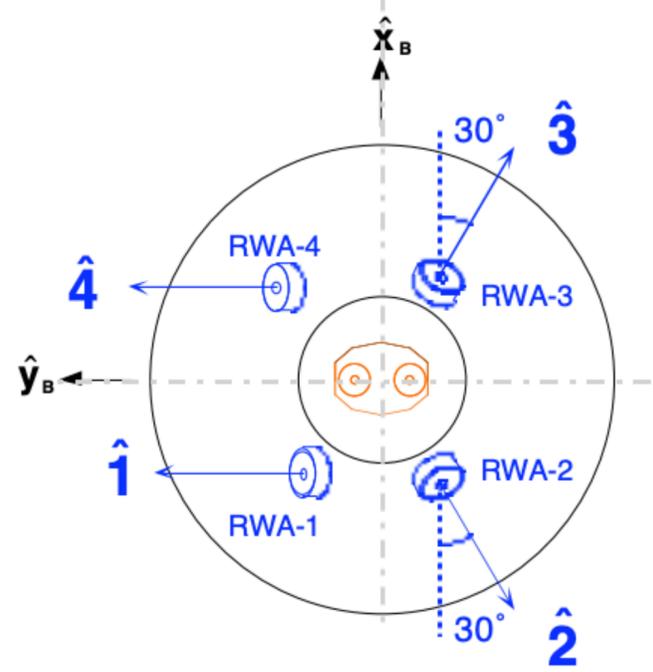
RWA Direction Cosines

$$\begin{bmatrix} \hat{1} \\ \hat{2} \\ \hat{3} \end{bmatrix} = \begin{bmatrix} 0 & \frac{\sqrt{2}}{\sqrt{3}} & \frac{1}{\sqrt{3}} \\ -\frac{1}{\sqrt{2}} & -\frac{1}{\sqrt{6}} & \frac{1}{\sqrt{3}} \\ \frac{1}{\sqrt{2}} & -\frac{1}{\sqrt{6}} & \frac{1}{\sqrt{3}} \end{bmatrix} \begin{bmatrix} \hat{x}_B \\ \hat{y}_B \\ \hat{z}_B \end{bmatrix}$$

RWA4 is Redundant, Articulate

$\hat{4}$ can be aligned with $\hat{1}$ or $\hat{2}$ or $\hat{3}$

At Launch, $\hat{4}$ & $\hat{1}$ are aligned



Reaction wheels

- Cassini carried 4 reaction wheels. 3 strapped down, the fourth on a movable platform
- Reaction wheels arranged to span R3

Reaction thrusters

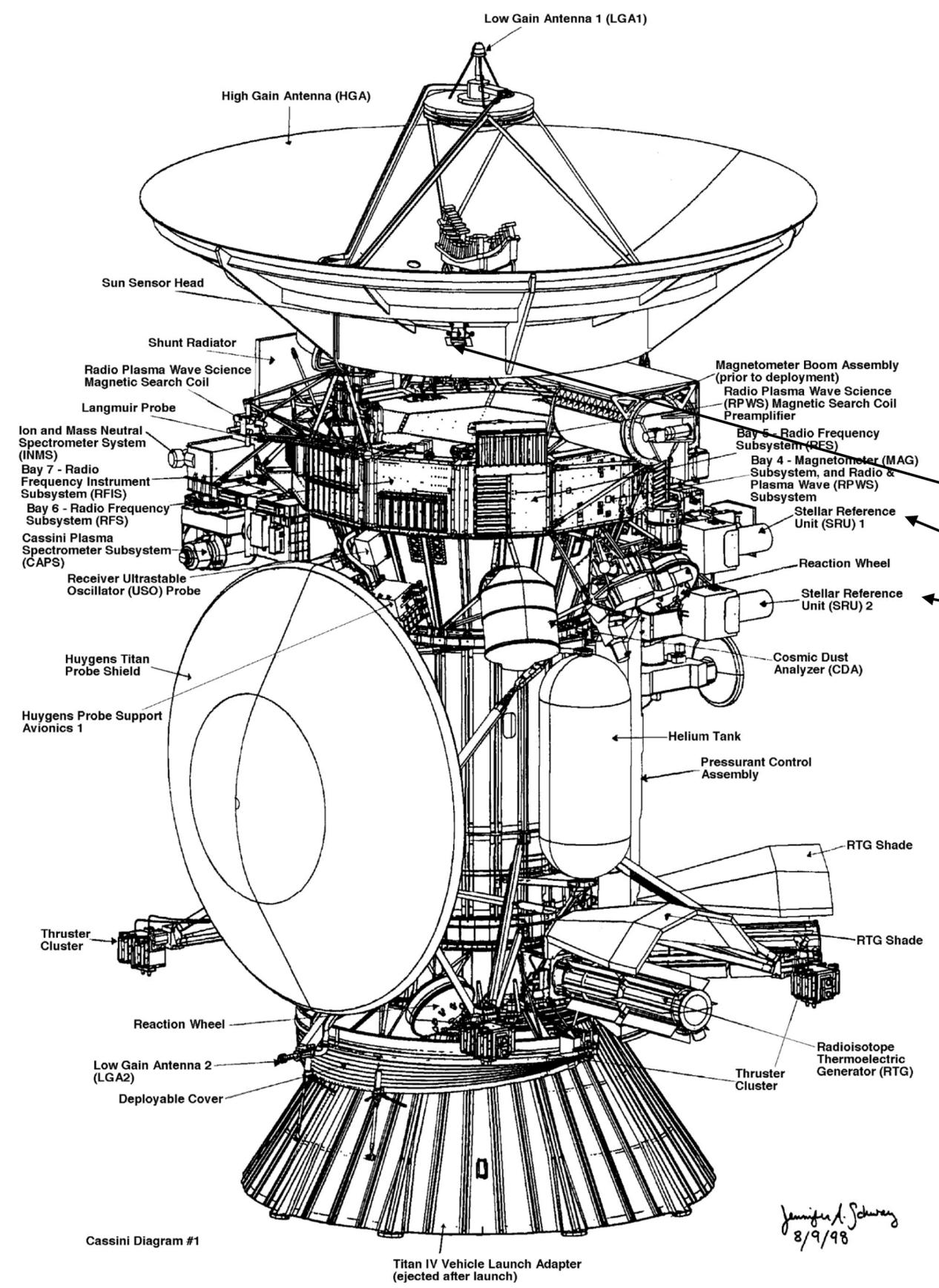
- 16 thrusters (8 prime, 8 redundant) for attitude control and momentum dumping.

Attitude Determination

What are the options?

What do you think Cassini used?

Attitude Determination



- Sun sensors (x2)
- Star trackers (x2)
- Fiber-optic gyros (x2)

Propulsion

- Delta-V capability of >2040 m/s (Though Cassini's total delta-V was much greater. **How?**)
- Two redundant main engines, fueled by N_2H_2/N_2O_4 (hydrazine and nitrogen tetroxide) bipropellant stored in conispherical tanks with capacity of 3450 kg
- Delta-V measured with accelerometer



Redundant engines

Guidance and Navigation



70 m DSN antenna in California

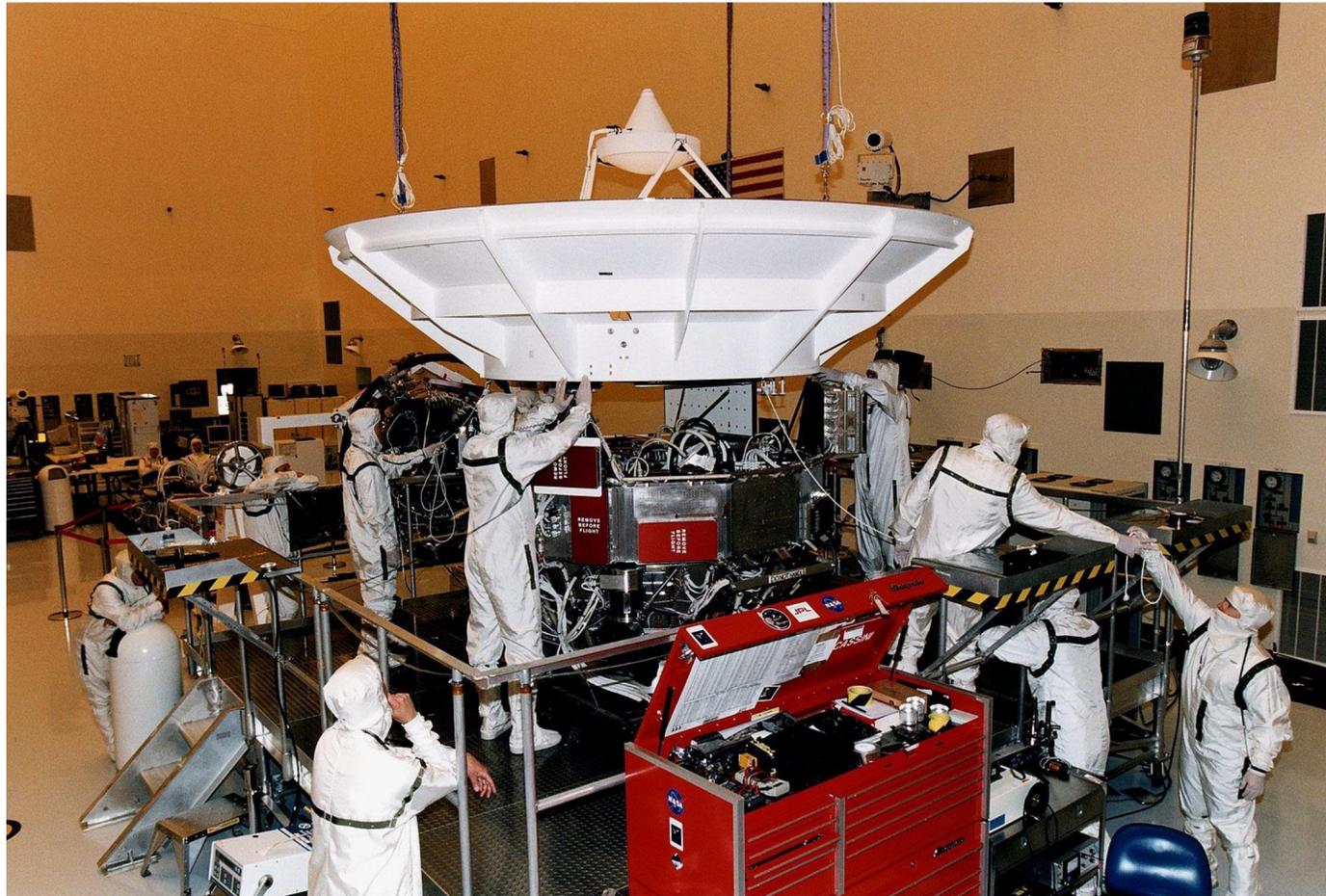
The Deep Space Network

- Network of communication facilities in California, Madrid, and Canberra.

Why these locations?

- Used for tracking deep-space satellites' positions and velocities

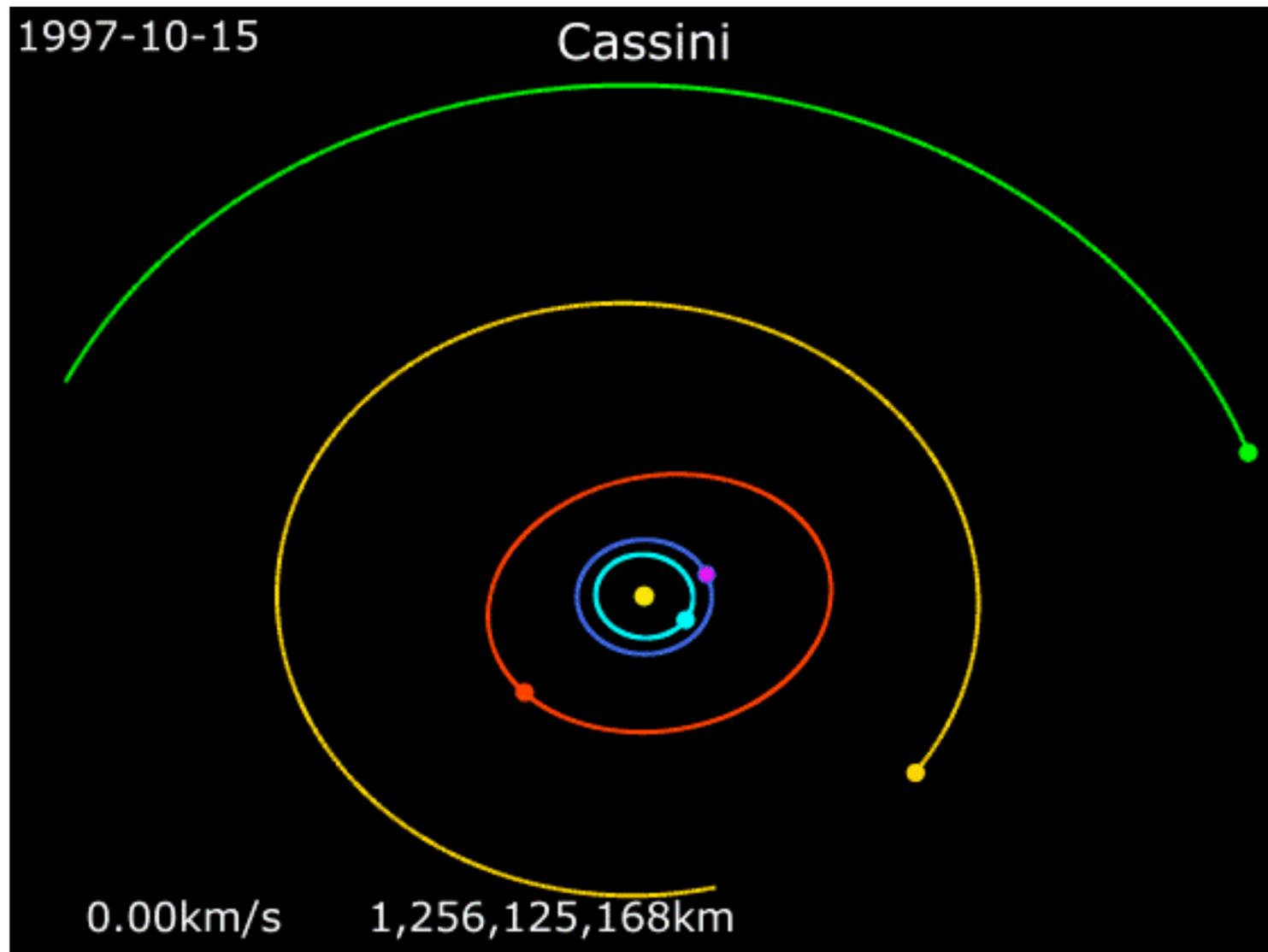
Telemetry and Command



High-gain antenna

- 3.67 meter high-gain X-band antenna (rebuild of *Voyager's*)
- 26.7 kbps from Saturn, with a bit error rate of $10e-6$ to 70m DSN antenna
- Reed-Solomon encoded
- Two low-gain X-band antennas for cruise phase and in the event of an emergency
- **How do we size these systems?**

Interplanetary Trajectory



Trajectory:

- Venus flyby 1 - Dec. 1998
- Venus flyby 2 - June 1999
- Earth flyby - August 1999
- Asteroid 2685 Masursky - Jan. 2000
- Jupiter flyby - Dec. 2000
- Saturn arrival - Feb. 2004

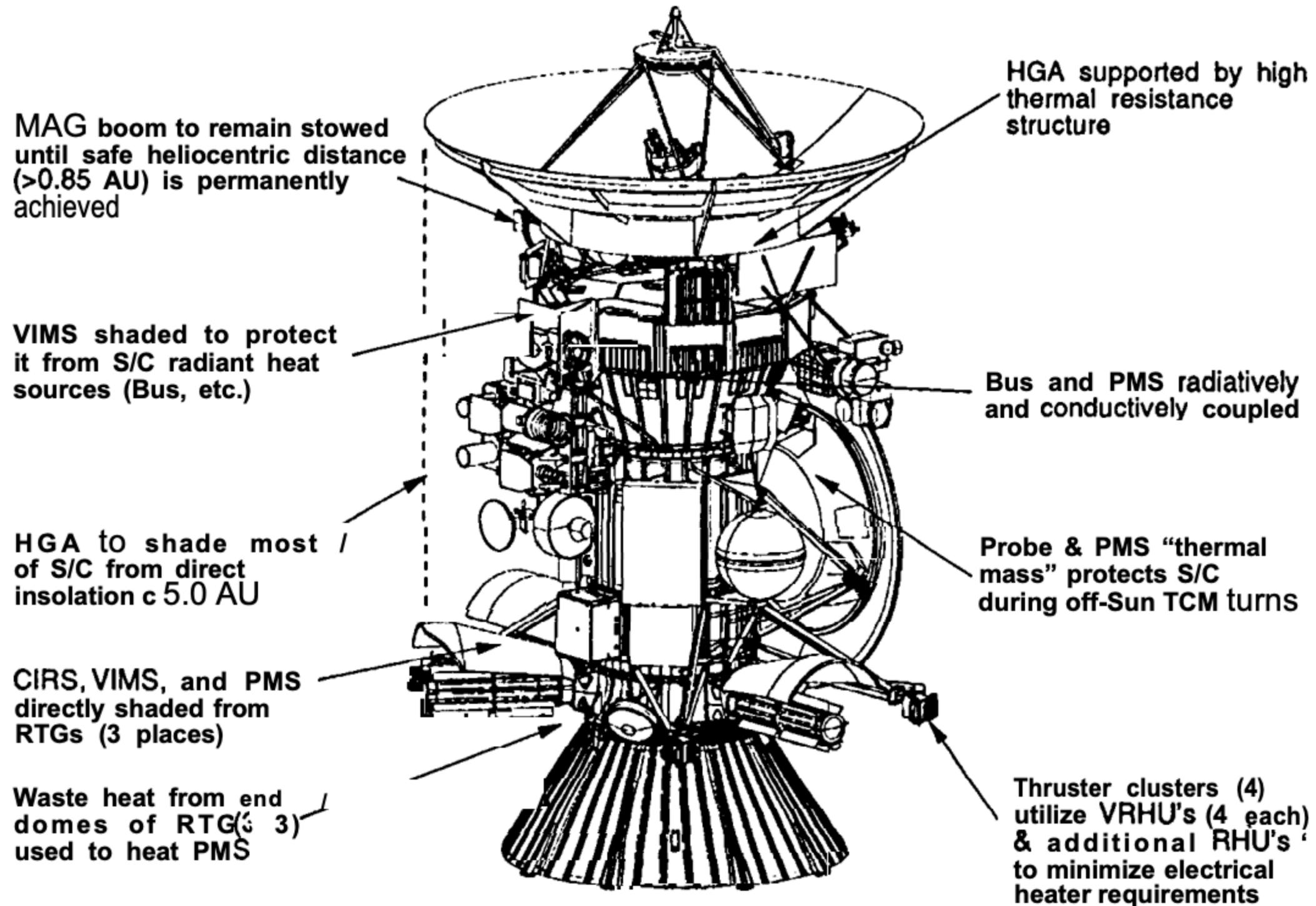
~0.86 - 9.2 AU

Thermal

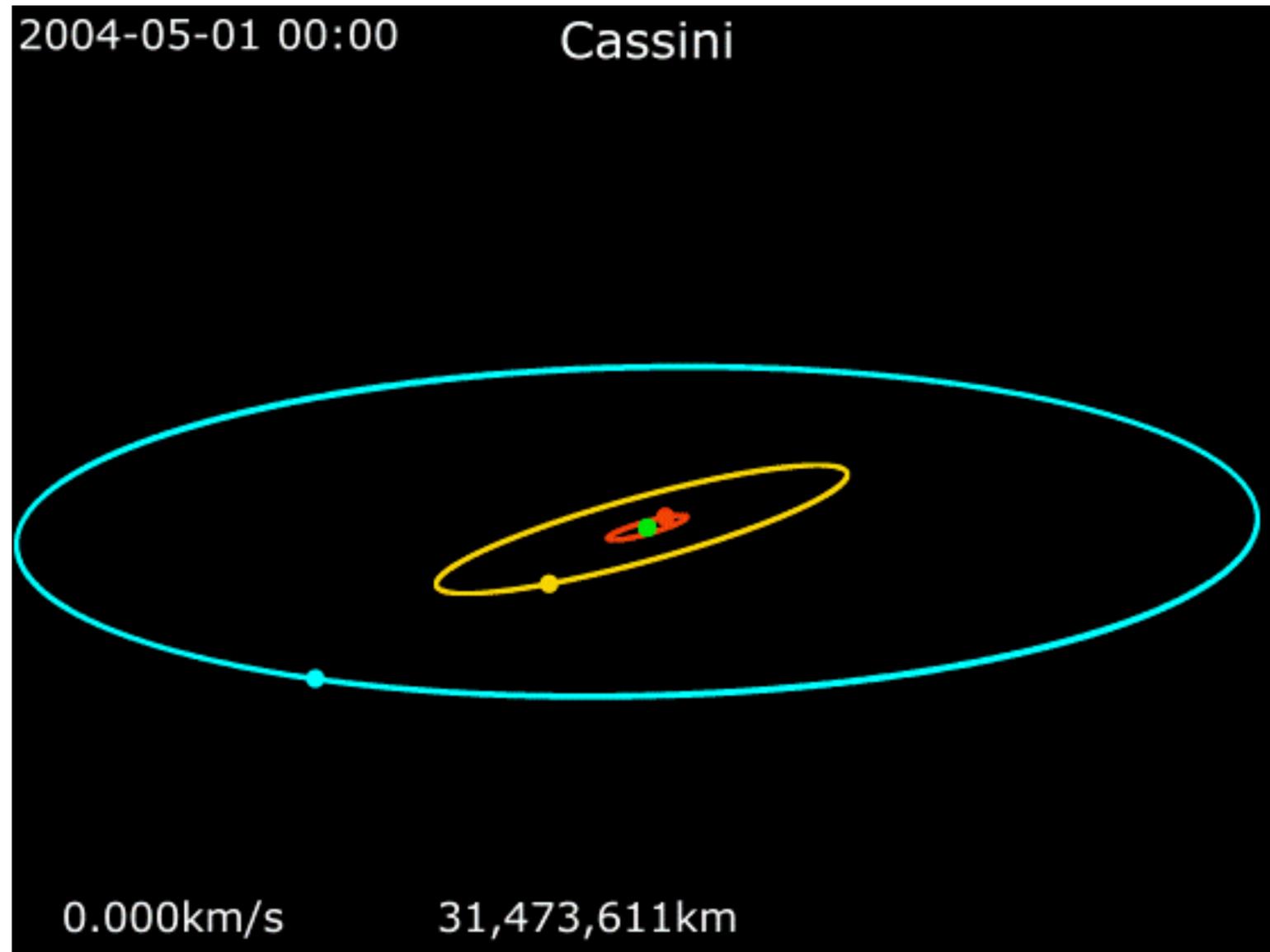
Why is this challenging, based on the trajectory that we just saw?

What other subsystem can we utilize for thermal management?

Thermal



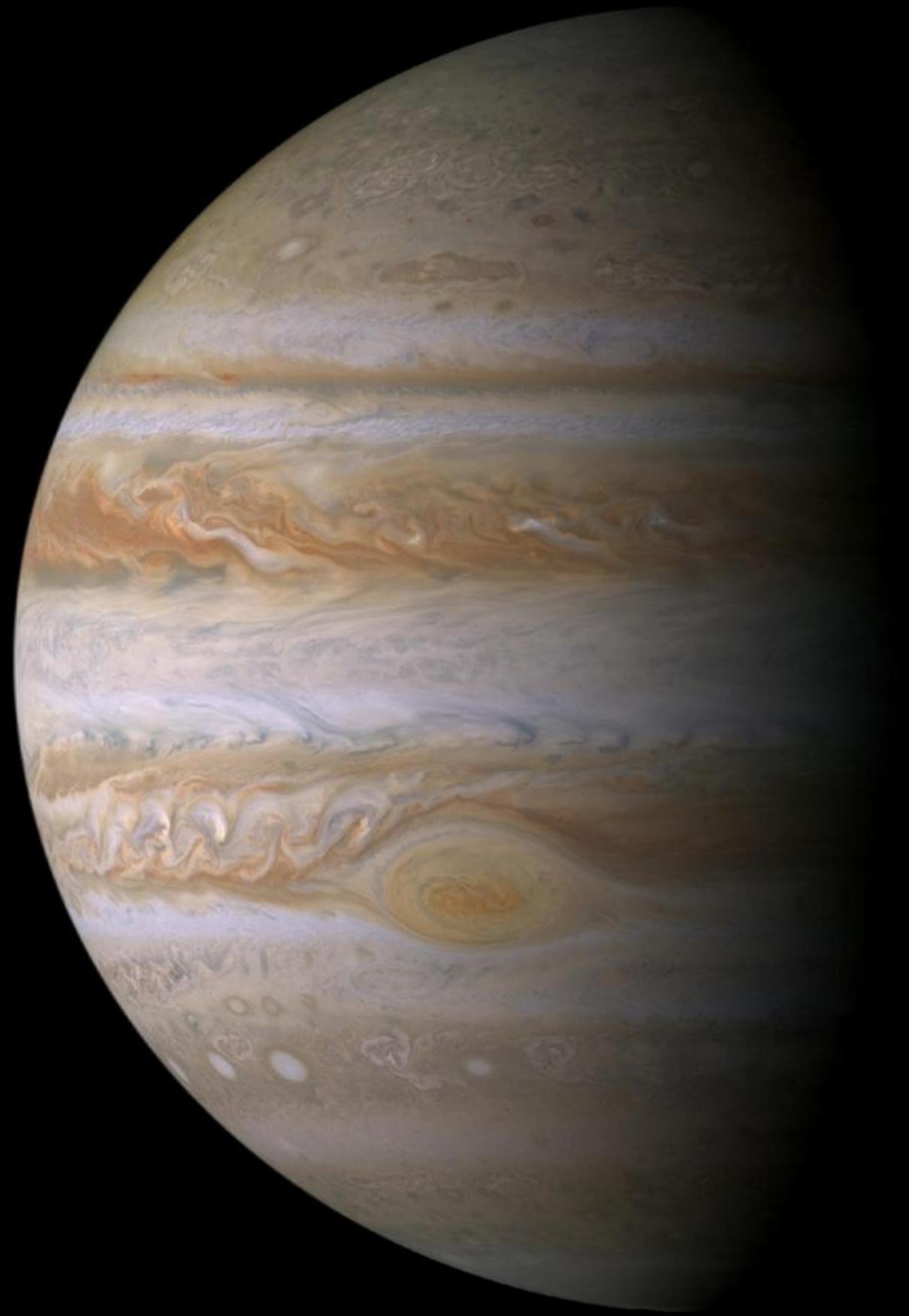
Saturn Trajectory

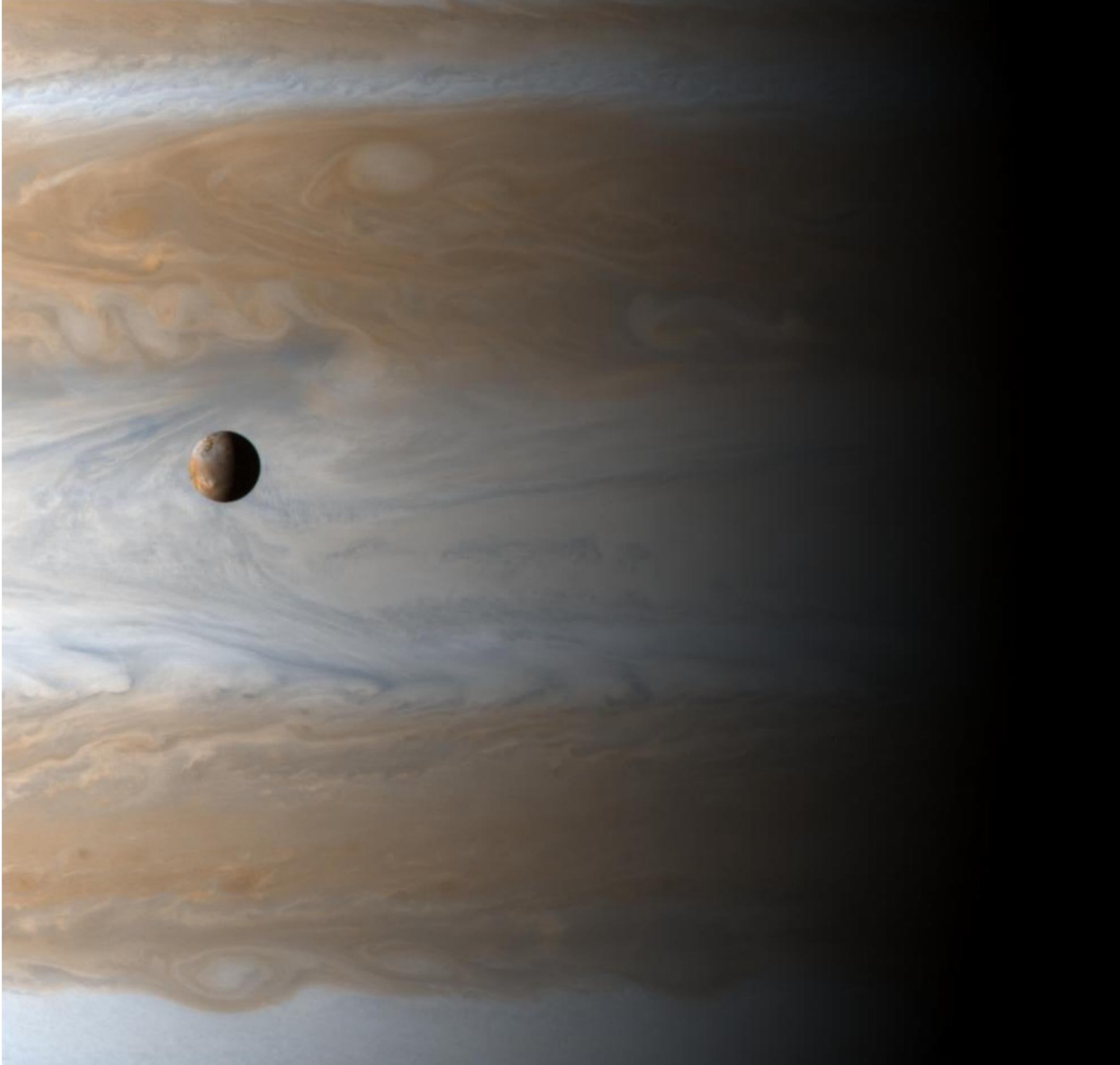


13 years at Saturn

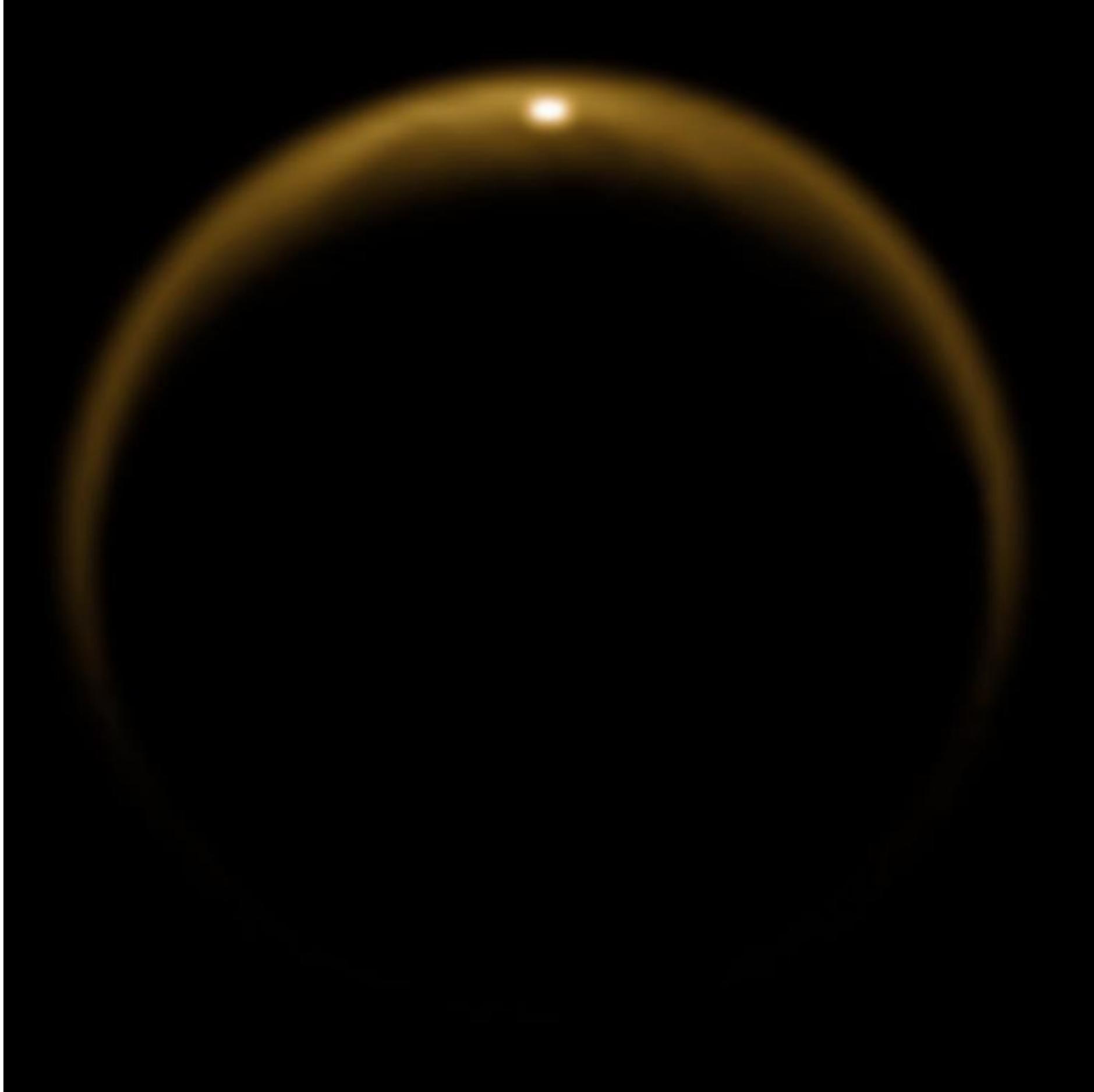
- Discovered lakes on Titan
- Observed a hurricane on Saturn
- Discovered 6 new moons
- Determined length of day on Saturn
- Imaged plumes of Enceladus
- ...

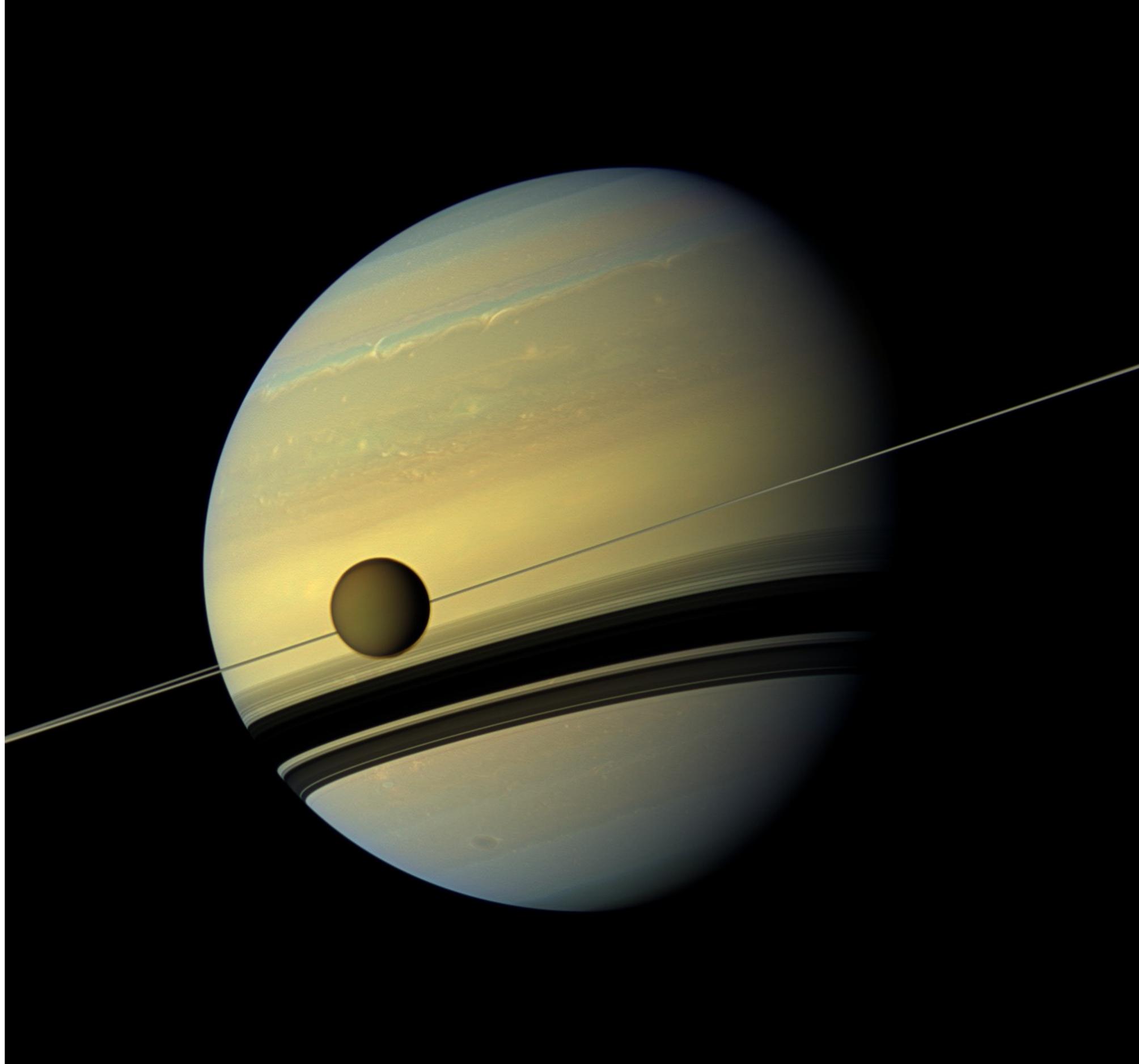
What did we get in return for this effort?

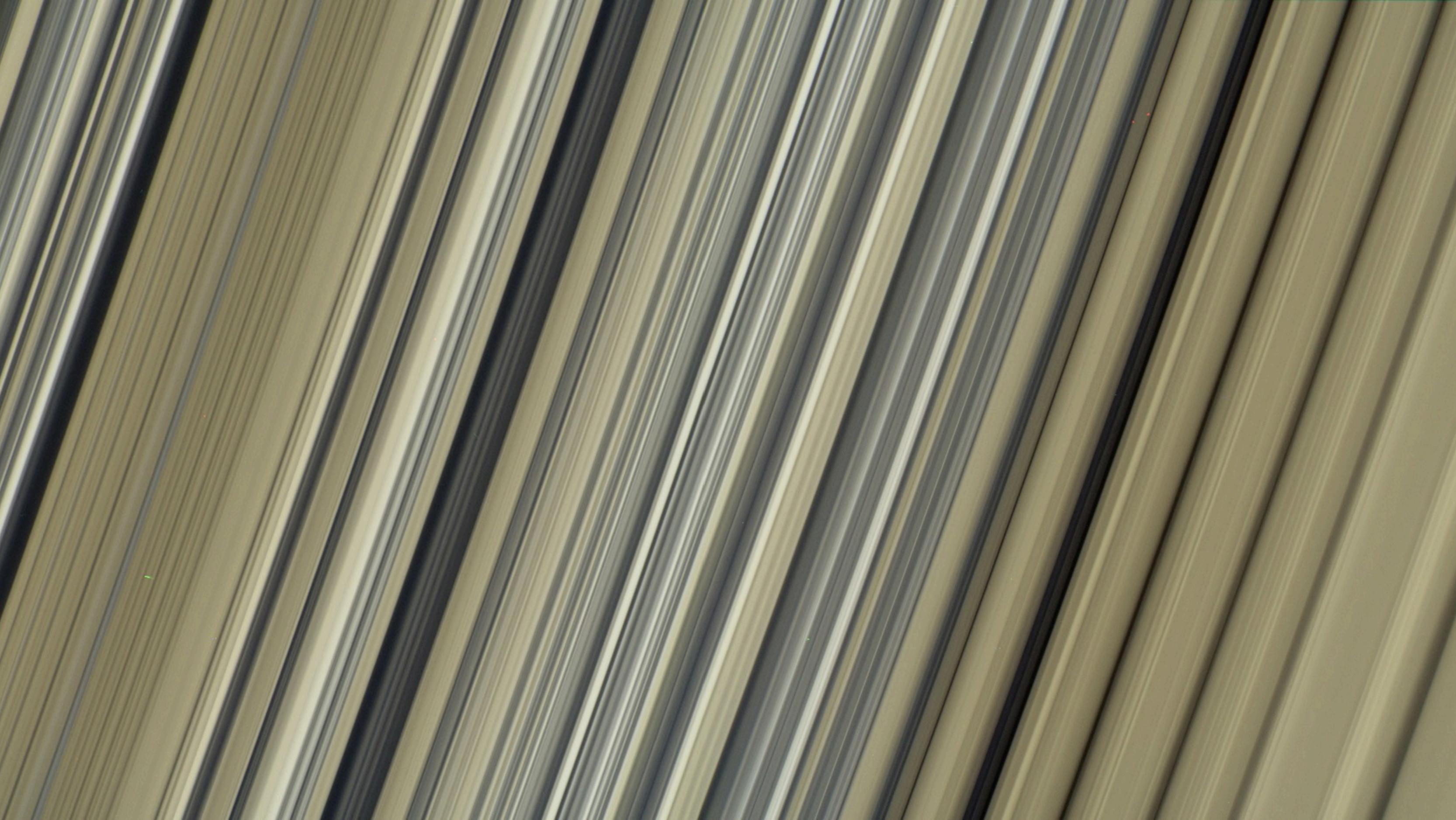


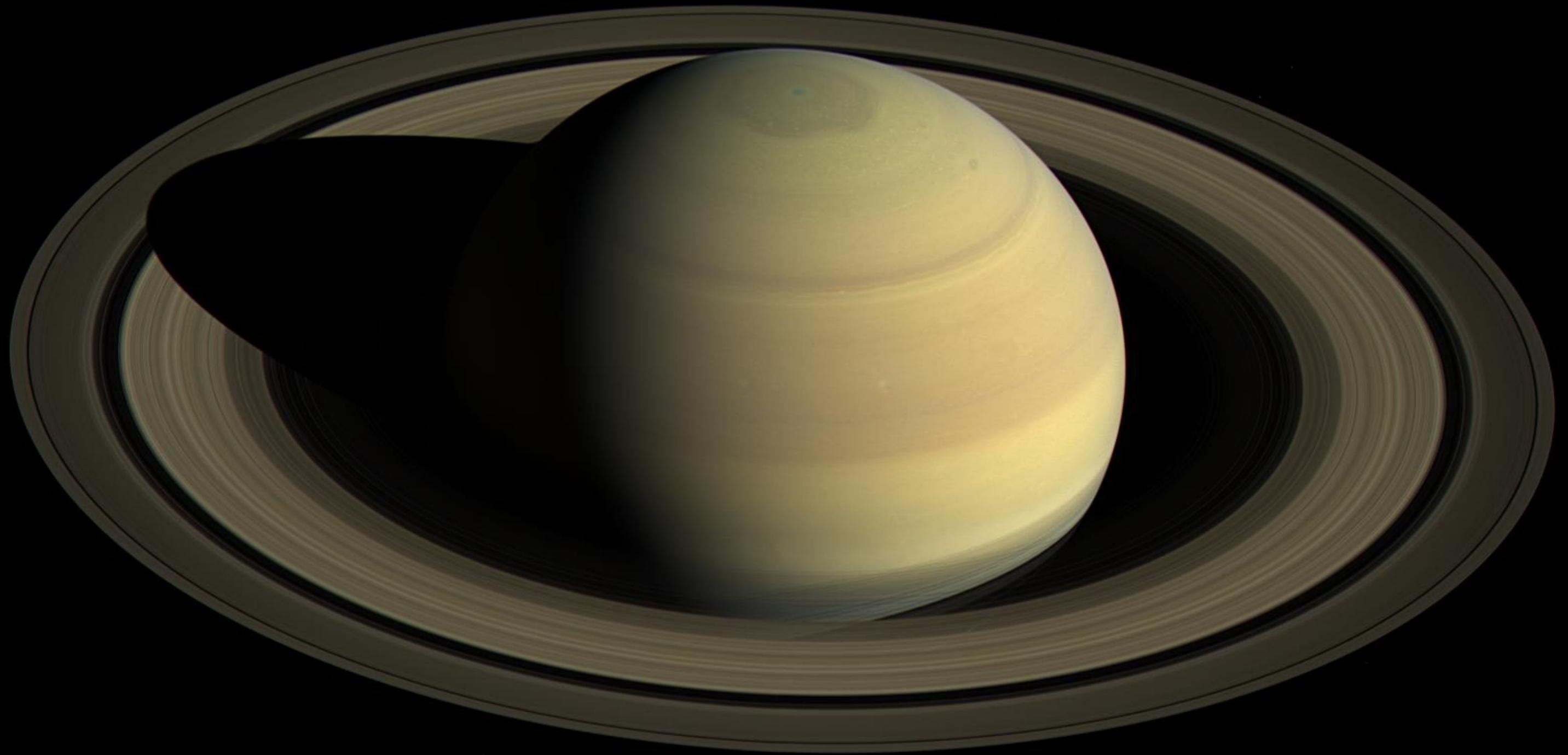


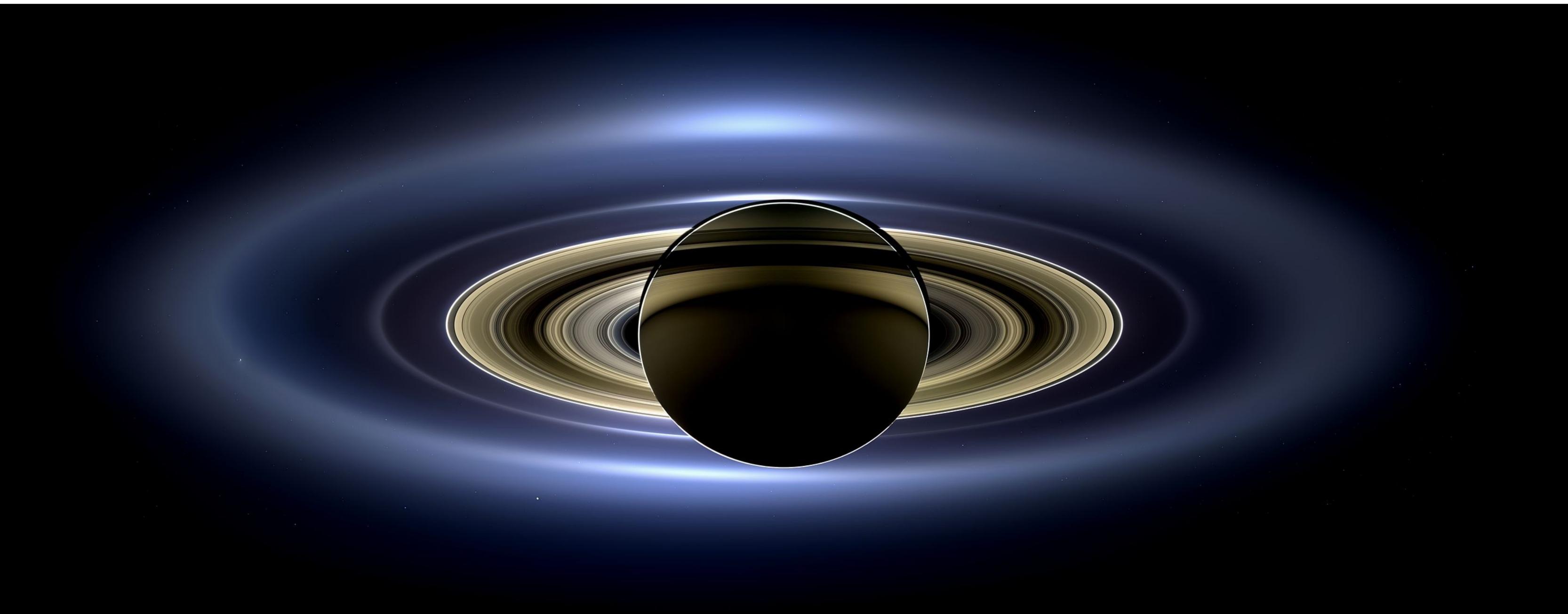


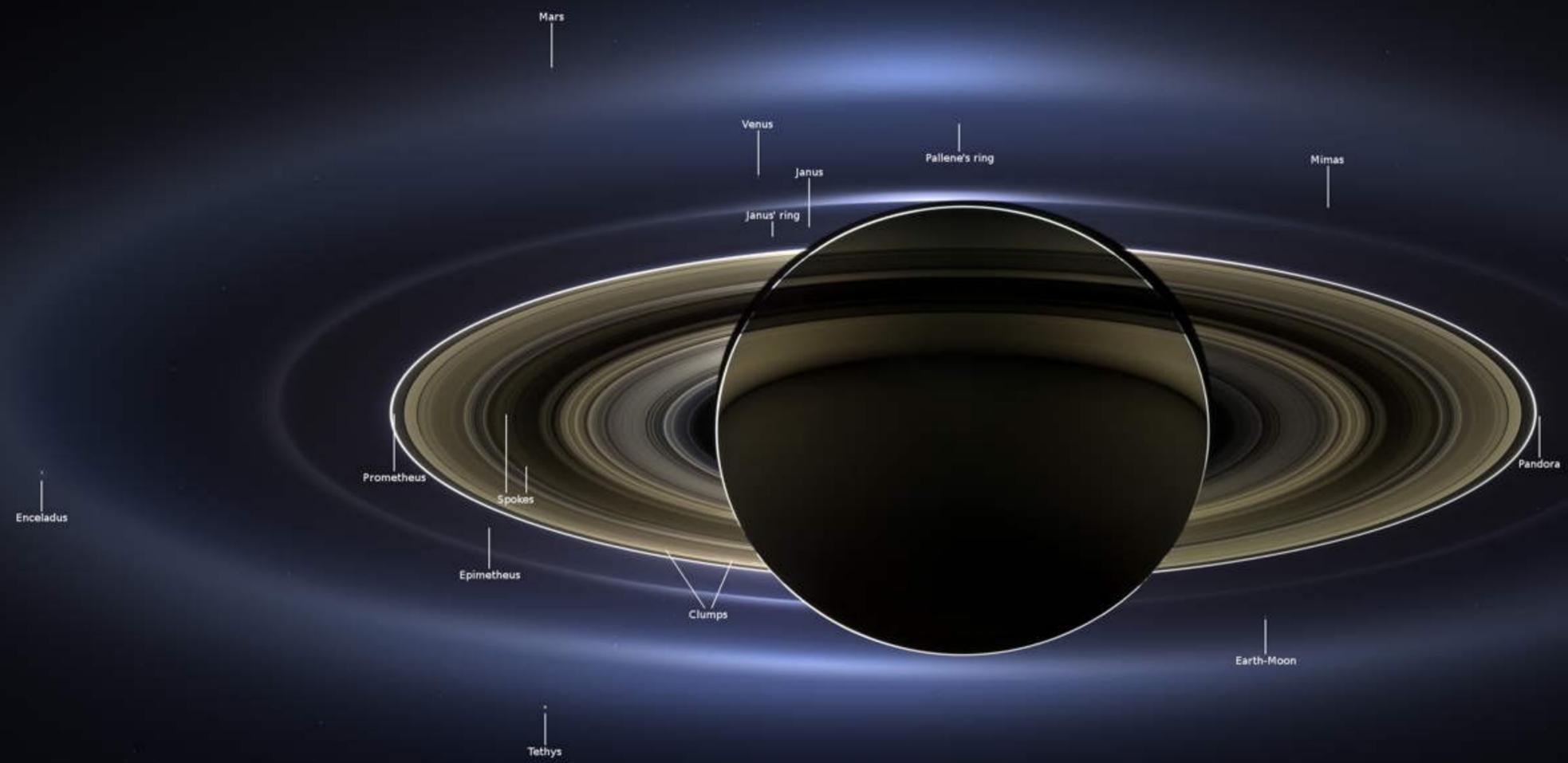




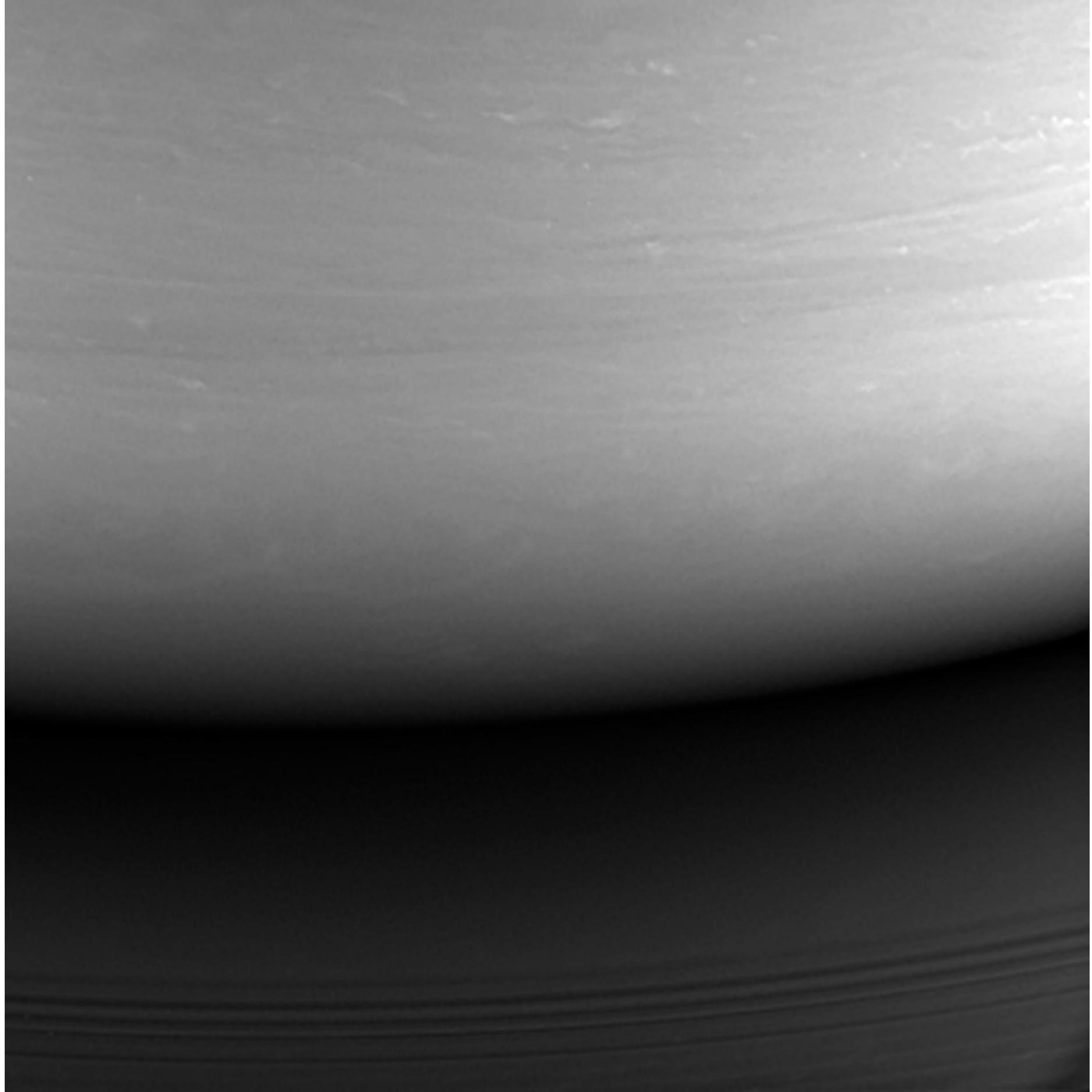








Why did we stop?



These are *places*. You can get there from here.

Before next time:

- Read the syllabus, come with questions
- Read lecture supplements 1 & 2